

Nonholonomic Systems, Dissipation and Quantization

Anthony M. Bloch

Work with Brockett, Ratiu, Marsden, Krishnaprasad, Zenkov, Hagerty, Weinstein, Rojo...

- Dissipation in nonholonomic systems
- Nonholonomic Systems and fields
- Quantization

- Class of systems exhibiting asymptotic stability behavior: nonholonomic systems – systems with nonintegrable constraints. In the absence of external dissipative forces, are always energy preserving.

Do not necessarily preserve volume in the phase space – – see for example Zenkov, Bloch and Marsden [1998], Zenkov and Bloch [2002], Kozlov, Jovanovich,

- Infinite Dimensions – oscillators interacting with fields. Hagerty, Bloch and Weinstein. Bloch, Hagerty, Rojo and Weinstein. Radiation Damping. Sofer and Weinstein.

- A nonholonomic system as a limit of system interacting with a field.

• **Double Brackets and Dissipation** Double bracket flows: dissipative mechanism in otherwise energy conserving mechanical systems, Bloch, Krishnaprasad, Marsden and Ratiu [1996].

• **Simple example: rigid body equations:**

$$I\dot{\Omega} = (I\Omega) \times \Omega,$$

or, in terms of the body angular momentum $M = I\Omega$,

$$\dot{M} = M \times \Omega.$$

Energy equals the Lagrangian: $E(\Omega) = L(\Omega)$ and energy is conserved.

Add a term cubic in the angular velocity:

$$\dot{M} = M \times \Omega + \alpha M \times (M \times \Omega),$$

where α is a positive constant.

- Related example is the Landau-Lifschitz equations for the magnetization vector M in a given magnetic field B :

$$\dot{M} = \gamma M \times B + \frac{\lambda}{\|M\|^2} (M \times (M \times B)),$$

where γ is the magneto-mechanical ratio (so that $\gamma\|B\|$ is the Larmour frequency) and λ is the damping coefficient due to domain walls.

- The equations are Hamiltonian with the rigid body Poisson bracket:

$$\{F, K\}_{\text{rb}}(M) = -M \cdot [\nabla F(M) \times \nabla K(M)]$$

with Hamiltonians given respectively by $H(M) = (M \cdot \Omega)/2$ and $H(M) = \gamma M \cdot B$.

Dissipation in these systems is not induced by *any* Rayleigh dissipation function in the *literal* sense

However, it is induced by a dissipation function in the following restricted sense: It is a gradient when restricted to each momentum sphere,

Have:

$$\frac{d}{dt} \|M\|^2 = 0$$

$$\frac{d}{dt} E = -\alpha \|M \times \Omega\|^2,$$

for the rigid body,

• Interesting feature of these dissipation terms is that they can be derived from a symmetric bracket. In much the same way that the Hamiltonian equations can be derived from a skew symmetric Poisson bracket. For the case of the rigid body, this bracket is

$$\{\{F, K\}\} = \alpha (M \times \nabla F) \cdot (M \times \nabla K).$$

(For more on symmetric brackets see Crouch [1981] and Lewis and Murray [1999].)

- The Chaplygin Sleigh

Here we describe the Chaplygin sleigh, perhaps the simplest mechanical system which illustrates the possible dissipative nature of energy preserving nonholonomic systems.

Nonholonomic: subject to nonintegrable constraints – satisfies Lagrange D'Alembert equations.

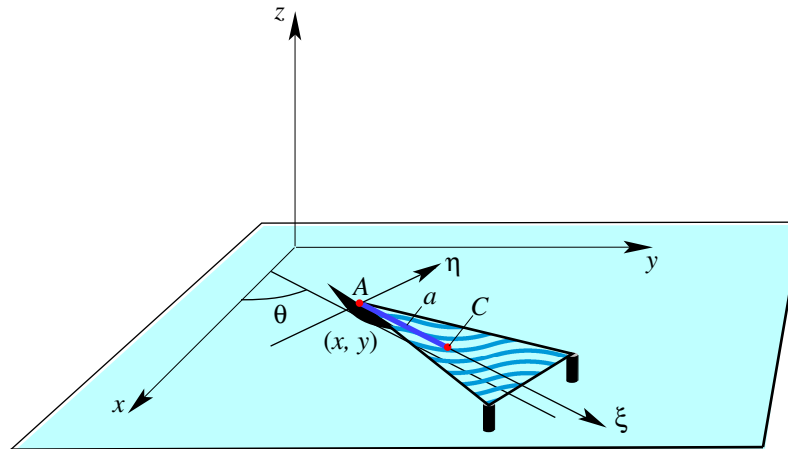


Figure 0.1: The Chaplygin sleigh is a rigid body moving on two sliding posts and one knife edge.

Equations:

$$\begin{aligned}\dot{v} &= a\omega^2 \\ \dot{\omega} &= -\frac{ma^2}{I+ma^2}v\omega\end{aligned}$$

**Equations have a family of relative equilibria given by $(v, \omega) | v =$
const, $\omega = 0$.**

Linearizing about any of these equilibria one finds one zero eigenvalue and one negative eigenvalue.

In fact the solution curves are ellipses in $v - \omega$ plane with the positive v -axis attracting all solutions.

Normalizing, we have the equations

$$\begin{aligned} \dot{v} &= \omega^2 \\ \dot{\omega} &= -v\omega. \end{aligned}$$

Scaling time by a factor of two have: Chaplygin sleigh equations are equivalent to the two-dimensional Toda lattice equations except for the fact that there is no sign restriction on the variable ω . Hence can be written in Lax pair form and solved by the method of factorization.

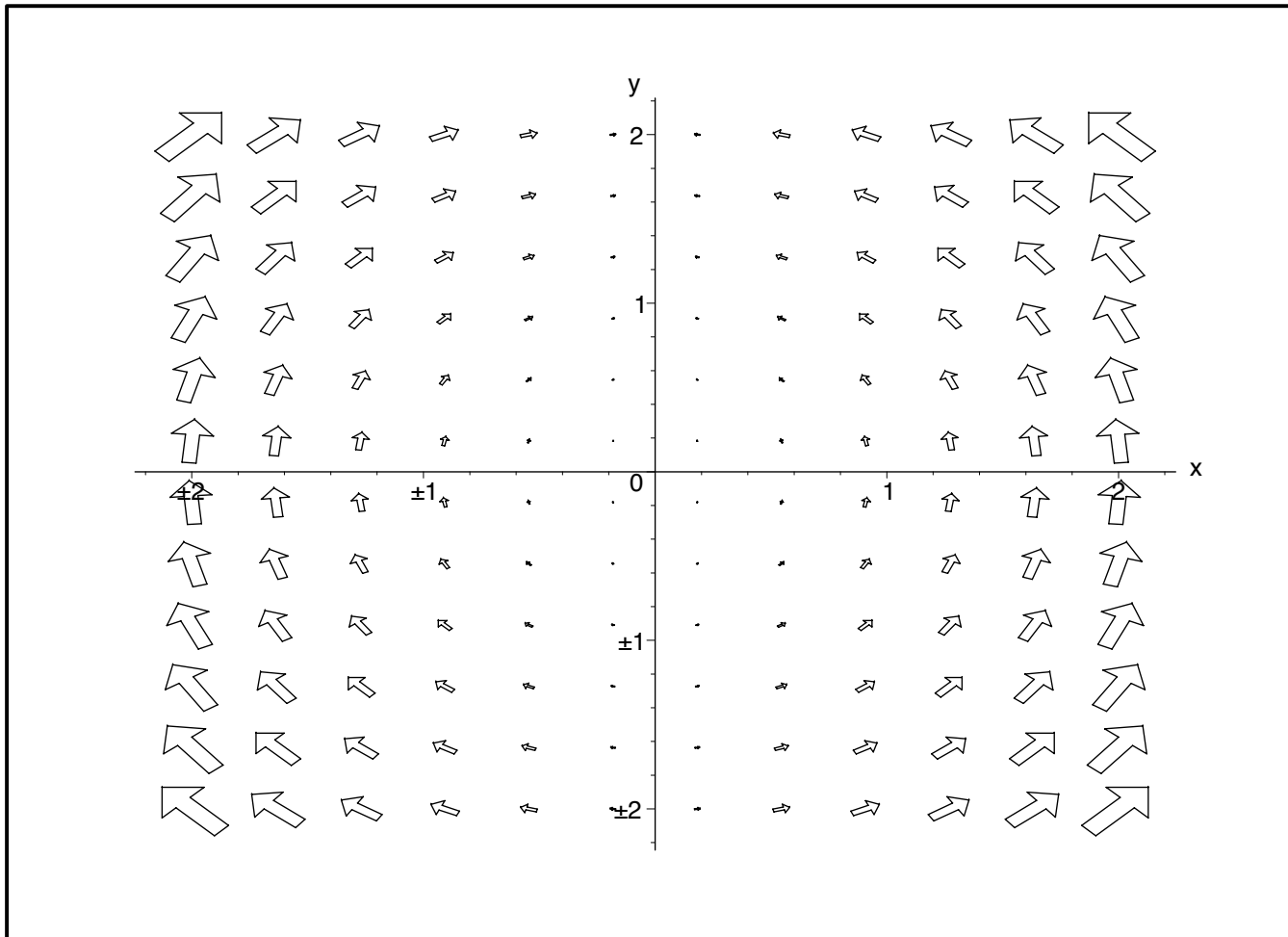


Figure 0.2: Chaplygin Sleigh/2d Toda phase portrait.

- Euler-Poincaré-Suslov Equations

Important special case of the reduced nonholonomic equations.

- Example: Euler-Poincaré-Suslov Problem on $SO(3)$ In this case the problem can be formulated as the standard Euler equations

$$I\dot{\omega} = I\omega \times \omega$$

where $\omega = (\omega_1, \omega_2, \omega_3)$ are the system angular velocities in a frame where the inertia matrix is of the form $I = \text{diag}(I_1, I_2, I_3)$ and the system is subject to the constraint

$$a \cdot \omega = 0$$

where $a = (a_1, a_2, a_3)$.

The nonholonomic equations of motion are then given by

$$I\dot{\omega} = I\omega \times \omega + \lambda a$$

subject to the constraint. Solve for λ :

$$\lambda = -\frac{I^{-1}a \cdot (I\omega \times \omega)}{I^{-1}a \cdot a}.$$

If a is an eigenvector of the moment of inertia tensor flow is measure preserving.

Invariant Measures of the Euler-Poincaré-Suslov Equations

An important special case of the reduced nonholonomic equations is the case when there is no shape space at all. In this case the system is characterized by the Lagrangian $L = \frac{1}{2}\mathbb{I}_{AB}\Omega^A\Omega^B$ and the left-invariant constraint

$$\langle a, \Omega \rangle = a_A \Omega^A = 0. \quad (0.1)$$

Here $a = a_A e^A \in \mathfrak{g}^*$ and $\Omega = \Omega^A e_A$, where e_A , $A = 1, \dots, k$, is a basis for \mathfrak{g} and e^A is its dual basis. Multiple constraints may be imposed as well. The two classical examples of such systems are the *Chaplygin Sleigh* and the *Suslov problem*. These problems were introduced by Chaplygin in 1895 and Suslov in 1902, respectively.

We can consider the problem of when such systems exhibit asymptotic behavior. Following Kozlov [1988] it is convenient to consider the unconstrained case first. In the absence of constraints the dynamics is governed by the basic Euler-Poincaré equations

$$\dot{p}_B = C_{AB}^C \mathbb{I}^{AD} p_C p_D = C_{AB}^C p_C \Omega^A \quad (0.2)$$

where $p_B = \mathbb{I}_{AB} \Omega^B$ are the components of the momentum $p \in \mathfrak{g}^*$. One considers the question of whether the (unconstrained) equations (0.2) have an absolutely continuous integral invariant $f d^k \Omega$ with summable density \mathcal{M} . If \mathcal{M} is a positive function of class C^1 one calls the integral invariant an invariant measure. Kozlov [1988] shows

Theorem 0.1 *The Euler-Poincaré equations have an invariant measure if and only if the group G is unimodular.*

A group is said to be unimodular if it has a bilaterally invariant measure. A criterion for unimodularity is $C_{AC}^C = 0$ (using the Einstein summation convention). Now we know (Liouville's theorem) that the flow of a vector differential equation $\dot{x} = f(x)$ is phase volume preserving if and only if $\text{Div } f = 0$. In this case the divergence of the right hand side of equation (0.2) is $C_{AC}^C \mathbb{I}^{AD} p_D = 0$. The statement of the theorem now follows from the following theorem of Kozlov [1998]: *A flow due to a homogeneous vector field in \mathbb{R}^n is measure-preserving if and only if this flow preserves the standard volume in \mathbb{R}^n .*

Now, turning to the case where we have the constraint (0.1) we obtain the *Euler-Poincaré-Suslov equations*

$$\dot{p}_B = C_{AB}^C \mathbb{I}^{AD} p_C p_D + \lambda a_B = C_{AB}^C p_C \Omega^A + \lambda a_B \quad (0.3)$$

together with the constraint (0.1). Here λ is the Lagrange multiplier. This defines a system on the subspace of the dual Lie algebra defined by the constraint. Since the constraint is assumed to be nonholonomic, this subspace is not a subalgebra. One can then formulate a condition for the existence of an invariant measure of the Euler-Poincaré-Suslov equations.

Theorem 0.2 *Equations (0.3) have an invariant measure if and only if*

$$K \operatorname{ad}_{\mathbb{I}^{-1}a}^* a + T = \mu a, \quad \mu \in \mathbb{R}, \quad (0.4)$$

where $K = 1/\langle a, \mathbb{I}^{-1}a \rangle$ and $T \in \mathfrak{g}^*$ is defined by $\langle T, \xi \rangle = \operatorname{Trace}(\operatorname{ad} \xi)$.

This theorem was proved by Kozlov [1988] for compact algebras and for arbitrary algebras by Jovanović [1998].

In coordinates, condition (0.4) becomes

$$KC_{AB}^C \mathbb{I}^{AD} a_C a_D + C_{BC}^C = \mu a_B.$$

For a compact algebra (0.4) becomes

$$[\mathbb{I}^{-1}a, a] = \mu a, \quad \mu \in \mathbb{R}, \quad (0.5)$$

where we identified \mathfrak{g}^* with \mathfrak{g} .

The proof of theorem 0.2 reduces to the computation of the divergence of the vector field in (0.3).

In the compact case only constraint vectors a which commute with $\mathbb{I}^{-1}a$ allow the measure to be preserved. This means that a and $\mathbb{I}^{-1}a$ must lie in the same maximal commuting subalgebra. In particular, if a is an eigenstate of the inertia tensor, the reduced phase volume is preserved. When the maximal commuting subalgebra is one-dimensional this is a necessary condition. This is the case for groups such as $SO(3)$.

Bloch and Zenkov extend these ideas to the case of internal variables.

- **Radiation Damping**

See Hagerty, Bloch and Weinstein [1999], [2002].

Important early work: Lamb [1900]. Related recent work may be found in Soffer and Weinstein [1998a,b] [1999] and Kirr and Weinstein [2001].

- Original Lamb model an oscillator is physically coupled to a string. The vibrations of the oscillator transmit waves into the string and are carried off to infinity. Hence the oscillator loses energy and is effectively damped by the string.

- **Lamb model**

$w(x, t)$ displacement of the string. with mass density ρ , tension T . Assuming a singular mass density at $x = 0$, we couple dynamics of an oscillator, q , of mass M :

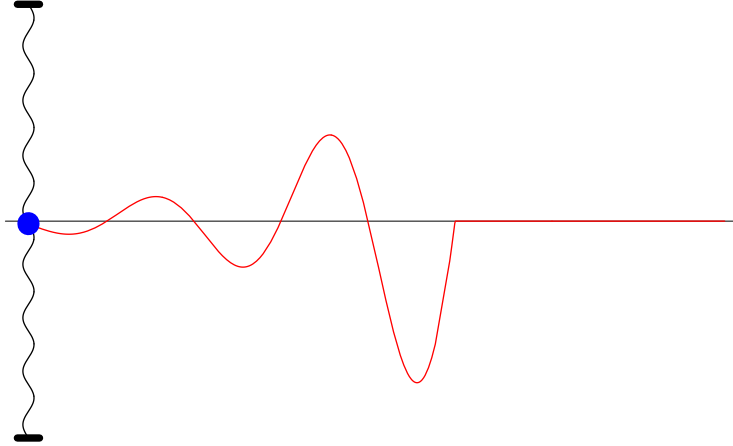


Figure 0.3: Lamb model of an oscillator coupled to a string.

$$\frac{\partial^2 w}{\partial t^2} = c^2 \frac{\partial^2 w}{\partial x^2}$$

$$M\ddot{q} + Vq = T[w_x]_{x=0}$$

$$q(t) = w(0, t).$$

$[w_x]_{x=0} = w_x(0+, t) - w_x(0-, t)$ is the jump discontinuity of the slope of the string. Note that this is a Hamiltonian system.

Can solve for w and reduce:

- Obtain a reduced form of the dynamics describing the explicit motion of the oscillator subsystem,

$$M\ddot{q} + \frac{2T}{c}\dot{q} + Vq = 0.$$

The coupling term arises explicitly as a Rayleigh dissipation term $\frac{2T}{c}\dot{q}$ in the dynamics of the oscillator.

Gyroscopic systems:

See Bloch, Krishnaprasad, Marsden and Ratiu [1994].

Linear systems of the form

$$M\ddot{q} + S\dot{q} + \Lambda q = 0$$

where $q \in \mathbb{R}^n$, M is a positive definite symmetric $n \times n$ matrix, S is skew, and Λ is symmetric.

This system Hamiltonian with $p = M\dot{q}$, energy function

$$H(q, p) = \frac{1}{2}pM^{-1}p + \frac{1}{2}q\Lambda q$$

and the bracket

$$\{F, K\} = \frac{\partial F}{\partial q^i} \frac{\partial K}{\partial p_i} - \frac{\partial K}{\partial q^i} \frac{\partial F}{\partial p_i} - S_{ij} \frac{\partial F}{\partial p_i} \frac{\partial K}{\partial p_j}.$$

Systems of this form arise from simple mechanical systems via reduction; normal form of the linearized equations when one has an *abelian* group.

Theorem 0.3 Dissipation induced instabilities—abelian case *Under the above conditions, if we modify the equation to*

$$M\ddot{q} + (S + \epsilon R)\dot{q} + \Lambda q = 0$$

for small $\epsilon > 0$, where R is symmetric and positive definite, then the perturbed linearized equations

$$\dot{z} = L_\epsilon z,$$

where $z = (q, p)$ are spectrally unstable, i.e., at least one pair of eigenvalues of L_ϵ is in the right half plane.

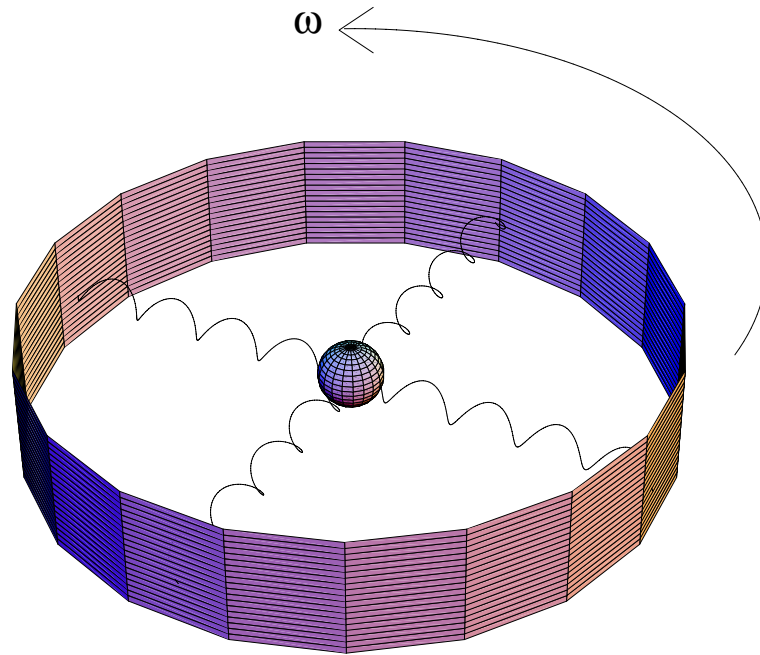


Figure 0.4: Rotating plate with springs.

- Gyroscopic systems connected to wave fields.

In Hagerty, Bloch and Weinstein [2002] we describe a gyroscopic version of the Lamb model coupled to a standard non-dispersive wave equation and to a dispersive wave equation. Show that instabilities will arise in certain mechanical systems.

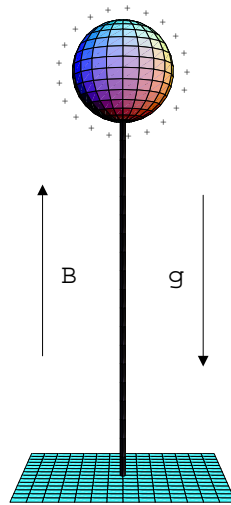


Figure 0.5: Inverted spherical pendulum.

In the dispersionless case, the system is of the form

$$\frac{\partial^2 \mathbf{w}}{\partial t^2}(z, t) = c^2 \frac{\partial^2 \mathbf{w}}{\partial z^2}(z, t),$$
$$M\ddot{\mathbf{q}}(t) + S\dot{\mathbf{q}}(t) + V\mathbf{q}(t) = T \left[\frac{\partial \mathbf{w}}{\partial z} \right]_{z=0}$$
$$\mathbf{w}(0, t) = \mathbf{q}(t),$$

$\mathbf{w} = [w_1(z, t) \cdots w_n(z, t)]^T$ is the displacement of the string in the first n dimensions and $[\frac{\partial \mathbf{w}}{\partial z}]_{z=0}$ is the jump discontinuity in the slope of the string.

- Can reduce dynamics to essentially:

$$M\ddot{\mathbf{q}}(t) = -S\dot{\mathbf{q}}(t) - V\mathbf{q}(t) - \frac{2T}{c}\dot{\mathbf{q}}(t),$$

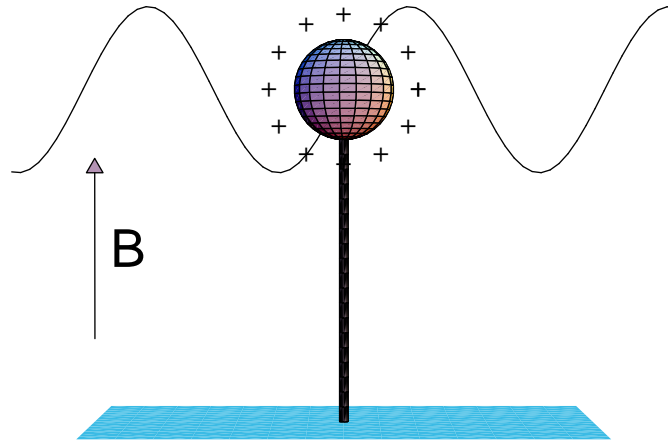


Figure 0.6: Gyroscopic Lamb coupling to a spherical pendulum.

- Non-local field coupling

$$M\ddot{\mathbf{q}} + S\dot{\mathbf{q}} + V\mathbf{q} = \kappa \int_{\mathbb{R}} \chi(z)w(z, t)dz \begin{bmatrix} 1 \\ \vdots \\ 1 \end{bmatrix},$$

$$\ddot{w} - c^2 \frac{\partial^2 w}{\partial z^2} = \kappa \chi(z) \begin{bmatrix} 1 \\ \vdots \\ 1 \end{bmatrix}^T \mathbf{q}$$

where κ is s coupling parameter and $\chi(\xi)$; is a suitable distribution.

- Nonholonomic Systems as Limits

It has been known (see even Cartheodory –1933) that the Lagrange–d’Alembert equations can be obtained by starting with an unconstrained system subject to appropriately chosen dissipative forces, and then letting these forces go to infinity in an appropriate manner.

Kozlov showed that the variational nonholonomic equations too can be obtained as the result of another limiting process: He added a parameter-dependent “inertial term” to the Lagrangian of the constrained system, and then showed that the unconstrained equations approach the variational equations as the parameter approaches infinity.

Nonholonomic constraints can be regarded in some sense as due to “infinite” friction. Several authors have asked if this can be quantified. Interestingly this goes back to the work at least of Caratheodory who asked if the limiting case of such friction could explain the motion of Chaplygin’s sleigh. Caratheodory claimed this could not be done but Fufaev (1964) showed that this was indeed possible. The general case was considered by Kozlov (1983) and Karapetyan (1983).

The key idea is to take a nonlinear Rayleigh dissipation function of the form

$$F = -\frac{1}{2}k \sum_{j=1}^m \left(\sum_{i=1}^n a_i^{(j)}(\mathbf{q}) \dot{q}_i \right)^2 \quad (0.6)$$

where $k > 0$ is a positive constant. Taking the limit as k goes to zero and using Tikhonov’s theorem yields the nonholonomic dynamics.

However, the system in this setting is still not Hamiltonian.

The goal here is to keep the system in the class of Hamiltonian systems by emulating the dissipation by coupling to an external field

- The Chaplygin Sleigh

This system consists of a rigid body moving on two sliding posts and one knife edge, and is perhaps the simplest n.h.s. containing the quasi-dissipative feature mentioned above. This mechanical system has three coordinates, two for the center of mass (x_C, y_C) and one “internal” angular variable θ for the rotation with respect to the knife edge located at $(x, y) = (x_C + a \cos \theta, y_C + a \sin \theta)$. The system can rotate freely around (x, y) but is only allowed to translate in the direction $(\cos \theta, \sin \theta)$: if we choose our coordinates as $\mathbf{q} = (x, y, \theta)$ there is a single constraint given by

$$\dot{x} \sin \theta - \dot{y} \cos \theta = 0, \quad (0.7)$$

or, $\mathbf{a}^{(1)} = (\sin \theta, -\cos \theta, 0)$.

The equations of motion can be obtained without resorting to the Lagrangian formalism, using simple balance of forces and can be expressed in the form:

$$\begin{aligned} \dot{v} &= a\omega^2, \\ \dot{\omega} &= -\frac{ma}{I + ma^2}v\omega, \end{aligned} \quad (0.8)$$

with $v = \dot{x} \cos \theta + \dot{y} \sin \theta$ the translational velocity, $\omega = \dot{\theta}$, m the mass and I the moment of inertia with respect to the center of mass. a is the distance between the center of mass and the contact point of the knife edge. The solutions of the above equations are ellipses in the (v, ω) plane with equilibria given by $\{v = \text{const}, \omega = 0\}$ and which are asymptotically stable.

The above equations can be also obtained using the virtual force method starting with the unconstrained Lagrangian

$$L_0 = \frac{m}{2} \left[\left(\dot{x} - a\dot{\theta} \sin \theta \right)^2 + \left(\dot{y} + a\dot{\theta} \cos \theta \right)^2 \right] + \frac{I}{2} \dot{\theta}^2, \quad (0.9)$$

and using a Lagrange multiplier in the equations of motion:

$$\begin{aligned} m \frac{d}{dt} \left(\dot{x} - a\dot{\theta} \sin \theta \right) &= -\lambda \sin \theta, \\ m \frac{d}{dt} \left(\dot{y} + a\dot{\theta} \cos \theta \right) &= \lambda \cos \theta, \\ (I + ma^2) \ddot{\theta} + ma\dot{\theta}(\dot{x} \cos \theta + \dot{y} \sin \theta) &= 0. \end{aligned} \quad (0.10)$$

Carathedory and Fufaev added a viscous friction force of the form

$$R = -Nu \quad (0.11)$$

to the sleigh equations, where u is the velocity in the direction perpendicular to the blade. (Note that interchange u and v compared to the original paper of Fufaev.)

Setting

$$k^2 = \frac{m}{I + ma^2}, \quad \epsilon = \frac{I}{Na^2} \quad (0.12)$$

the equations with dissipation become

$$u = \epsilon a \dot{\omega} \quad (0.13)$$

$$\dot{v} = a\omega^2 + \epsilon a \omega \dot{\omega} \quad (0.14)$$

$$ak^2 \dot{\omega} + v\omega = -\epsilon a \ddot{\omega} \quad (0.15)$$

It is clear that as ϵ goes to zero one recovers the original equations. Cartheodory incorrectly argued however that since no matter how small ϵ is these equations yield trajectories which

differ from that of the original system, dissipation cannot yield the nonholonomic constraints.

Fufaev realized this is not correct since the system degenerates from a system of three to two equations and thus there is a singularity. Setting $\mu = \epsilon a$ and $\sigma = \dot{\omega}$ we then get

$$\dot{\omega} = \sigma \tag{0.16}$$

$$\dot{v} = a\omega^2 + \mu\omega\sigma \tag{0.17}$$

$$\mu\dot{\sigma} = -ak^2\sigma - v\omega \tag{0.18}$$

Then as $\mu \rightarrow 0$ we get rapid motion except for the surface

$$ak^2\sigma + \mu\omega = 0. \tag{0.19}$$

The slow motion of this surface onto the v - ω plane then gives the correct equations of motion.

- The Chaplygin Sleigh as a Particle in a Radiation Field

We now show that the sleigh equations can be obtained from a variational principle as reduced equations of motion after the system is coupled to an environment described by an $U(1)$ infinite field of the form $\mathbf{a}(\mathbf{z}, t) \equiv [\cos \alpha(\mathbf{z}, t), \sin \alpha(\mathbf{z}, t)]$. For the Lagrangian of the free field we choose

$$L_F = \frac{K}{2} \int d^2\mathbf{z} \dot{\mathbf{a}}^2, \quad (0.20)$$

and we couple the sleigh and the field with a term of the form

$$L_1 = \int d^2\mathbf{z} \delta(\mathbf{z} - \mathbf{x}) [\gamma \dot{\mathbf{x}} \cdot \mathbf{a} + \mu \cos(\alpha(\mathbf{z}, t) - \theta)]. \quad (0.21)$$

The first term in square brackets corresponds to a minimal coupling that favors $\dot{\mathbf{x}}$ in the direction of \mathbf{a} ; the second has the form of a potential coupling that favors an alignment of the internal variable θ with the local direction of \mathbf{a} .

The total action is $S = \int dt(L_0 + L_F + L_1)$ where L_0 is the Lagrangian of the free sleigh

$$L_0 = \frac{m}{2} \left[\left(\dot{x} - a\dot{\theta} \sin \theta \right)^2 + \left(\dot{y} + a\dot{\theta} \cos \theta \right)^2 \right] + \frac{I}{2} \dot{\theta}^2, \quad (0.22)$$

and can be regarded as a full “microscopic” theory of the sleigh coupled to an environment.

The equations of motion of the combined system are now obtained from a variational principle, $\delta S = 0$, and have the form

$$\begin{aligned}
& m \frac{d}{dt} \left(\dot{x} - a\dot{\theta} \sin \theta \right) \\
= & \gamma \left\{ -\sin \alpha(\mathbf{x}, t) \frac{\partial \alpha}{\partial t} + [\dot{x} \sin \alpha(\mathbf{x}, t) - \dot{y} \cos \alpha(\mathbf{x}, t)] \frac{\partial \alpha}{\partial x} \right\} \\
& - \mu \sin[\alpha(\mathbf{x}, t) - \theta] \frac{\partial \alpha}{\partial x}, \\
& m \frac{d}{dt} \left(\dot{y} + a\dot{\theta} \cos \theta \right) \\
= & \gamma \left\{ \cos \alpha(\mathbf{x}, t) \frac{\partial \alpha}{\partial t} + [\dot{x} \sin \alpha(\mathbf{x}, t) - \dot{y} \cos \alpha(\mathbf{x}, t)] \frac{\partial \alpha}{\partial y} \right\} \\
& - \mu \sin[\alpha(\mathbf{x}, t) - \theta] \frac{\partial \alpha}{\partial y}, \\
& (I + ma^2)\ddot{\theta} - ma \frac{d}{dt} (\dot{x} \sin \theta - \dot{y} \cos \theta) - ma\dot{\theta} (\dot{x} \cos \theta + \dot{y} \sin \theta) \\
= & \mu \sin [\alpha(\mathbf{x}, t) - \theta], \\
& K \frac{\partial^2 \alpha(\mathbf{z}, t)}{\partial t^2} \\
= & \delta(\mathbf{z} - \mathbf{x}) \{ \gamma [\dot{x} \sin \alpha(\mathbf{x}, t) - \dot{y} \cos \alpha(\mathbf{x}, t)] + \mu \sin [\alpha(\mathbf{x}, t) - \theta] \} (0.23)
\end{aligned}$$

At this point we take the limit $\mu \rightarrow \infty$ in the third equation above. This limit can be understood from the singular perturbation theory, by dividing the left hand side of the equation by μ , which amounts to rescaling the times in the derivatives by $\sqrt{\mu}$. (This is immediate by noting that the r.h.s. is homogeneous in the derivatives.) Therefore, for very large μ we have a very slow dynamics on the r.h.s., which amounts to setting $\sin[\alpha(\mathbf{x}, t) - \theta] = 0$. This is equivalent to saying that in the $\mu \rightarrow \infty$ limit the variables $\alpha(\mathbf{x}, t)$ and θ are pinned to the same value. Next we integrate equation (0.23) over an infinitesimal region around \mathbf{x} and obtain

$$\dot{x} \sin \alpha(\mathbf{x}, t) - \dot{y} \cos \alpha(\mathbf{x}, t) = \dot{x} \sin \theta - \dot{y} \cos \theta = 0, \quad (0.24)$$

which means that the constraint is satisfied. Replacing the constraint (and $\sin[\alpha(\mathbf{x}, t) - \theta] = 0$) in the first three equations we obtain the same structure as (0.10) and therefore the same flow as in Eq. (0.8).

The calculation shows that we have succeeded in deriving the nonholonomic equations for a system with one internal (compact) variable from a pure Lagrangian formalism. The classical trajectories are obtained from a variational principle and quantization can be introduced through the Path integral formalism: the propagator is $e^{iS/\hbar}$, where S is the complete action. Also, for the case of the sleigh, and for compact internal variables, we can give an intuitive description of the classical evolution. The sleigh is coupled to an infinite bath of rotors and, for $\mu \rightarrow \infty$, the internal variable and the rotors are locally the same. In the limit $K \rightarrow 0$ (vanishing moment of inertia for the rotors) the internal variable imposes its value on the local field instantaneously. Since the rotors are fixed in space they can still guide the motion imposing the velocity to be locally parallel to \mathbf{a} . Also, since we are taking the limit $K \rightarrow 0$, the field does not take energy from the sleigh, and the nonholonomic motion conserves energy.

Quantum Field Theory

Quantum case for $\mathbf{a}=0$:

The Hamiltonian in this limit has the form

$$H = \frac{1}{2m} [p_x - \lambda \cos \alpha(\mathbf{x})]^2 + \frac{1}{2m} [p_y - \lambda \sin \alpha(\mathbf{x})]^2 + \frac{1}{2I} p_\theta^2 \quad (0.25)$$

$$+ \frac{1}{2K} \int d\mathbf{z} \Pi^2(\alpha(\mathbf{z})) + \mu \cos[\theta - \alpha(\mathbf{x})]. \quad (0.26)$$

For the quantization of H we proceed with the usual replacements

$$\mathbf{p} = -i\hbar(\partial_x, \partial_y), \quad p_\theta = -i\hbar\partial_\theta, \quad \Pi(\alpha(\mathbf{z})) = -i\hbar\partial_{\alpha(\mathbf{z})}. \quad (0.27)$$

For the completely uncoupled case ($\lambda = \mu = 0$) the eigenstates are of the form

$$\Psi_0 = e^{i \int d\mathbf{z} m(\mathbf{z})\alpha(\mathbf{z})} e^{i\mathbf{k}\cdot\mathbf{x}} e^{in\theta}, \quad (0.28)$$

with $m(\mathbf{z})$ and n integers and \mathbf{k} the wave number of the translational degree of freedom.

The limit $\mu \rightarrow \infty$ amounts to projecting the wave function and the Hamiltonian to states where $\alpha(\mathbf{x}) = \theta$, in such a way that the Hamiltonian becomes

$$H = \frac{1}{2m} [p_x - \lambda \cos \theta]^2 + \frac{1}{2m} [p_y - \lambda \sin \theta]^2 + \frac{1}{2I'} p_\theta^2 \quad (0.29)$$

$$+ \frac{1}{2K} \int d\mathbf{z} \Pi^2(\alpha(\mathbf{z})) [1 - \delta(\mathbf{x} - \mathbf{z})], \quad (0.30)$$

with $1/I' = 1/I + 1/K$. Without loss of generality we can take the quantum numbers $m[(\mathbf{z}) = 0$ for $\mathbf{z} \neq \mathbf{x}$ and the wave function depends only on the $\{\theta, \mathbf{x}\}$ degrees of freedom and obeys the following Schoedinger equation:

$$\left\{ \frac{1}{2m} [p_x - \lambda \cos \theta]^2 + \frac{1}{2m} [p_y - \lambda \sin \theta]^2 + \frac{1}{2I'} p_\theta^2 \right\} \Psi = \epsilon \Psi. \quad (0.31)$$

The above equation can be solved by separation of variables $\Psi = e^{i\mathbf{k}\cdot\mathbf{x}} \psi_{\mathbf{k}}(\theta)$, with $\mathbf{k} = k(\cos \theta_{\mathbf{k}}, \sin \theta_{\mathbf{k}})$ a quasi-translational wave-

vector. The reduced equation satisfied the the angular part of the wave function is

$$\left\{ \frac{1}{2I'} p_\theta^2 - \frac{\lambda \hbar k}{m} \cos(\theta - \theta_{\mathbf{k}}) \right\} \psi_{\mathbf{k}}(\theta) = \epsilon' \psi_{\mathbf{k}}(\theta), \quad (0.32)$$

with $\epsilon' = \epsilon - (\lambda^2 + \hbar^2 k^2)/2m$. This equation has well known solutions in terms of the Mathieu functions. One can gain insight on the structure of the solutions by looking at the fast limit ($k \rightarrow \infty$) which should exhibit features of the classical solution. In this limit the fluctuations of the angle are small and centered around $\theta = \theta_{\mathbf{k}}$. This means that, up to small quantum fluctuations, the knife edge is pointing in the direction of the plane wave propagation. Expanding for small values of the angle we find that the solutions in the fast limit are of the form

$$\Psi_{\mathbf{k}}(\mathbf{x}, \theta) = e^{i\mathbf{k} \cdot \mathbf{x}} e^{-(\theta - \theta_{\mathbf{k}})^2 / 2\Delta_\theta^2}, \quad (0.33)$$

with

$$\Delta_\theta^2 = \frac{m\hbar}{\lambda k I'}. \quad (0.34)$$

Some key refs: papers written by the author in collaboration with:

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